

## **EAS Maximum Depth from the Space-Time Structure of Cherenkov Light Based on the TAIGA-HiSCORE Data**

A new simulation using the CORSIKA code without accelerating coarsening of the results was performed for the energy range of 1-100 PeV, taking into account pulse formation at the input of the optical station of the TAIGA-HiSCORE array. Refined relationships between the steepness parameter of the lateral distribution function (LDF) and the light pulse FWHM at distances of 200-400 m and the relative position of the EAS maximum have been obtained. The application of the new simulated relationships to measure the EAS maximum depth ( $X_{max}$ ) by the data of the 3 seasons of the TAIGA-HiSCORE observations (2021 – 2024) are shown.

**Primary author:** TERNOVOY, Mark (ISU)

**Co-author:** TAIGA COLLABORATION

**Presenter:** TERNOVOY, Mark (ISU)

**Session Classification:** Cosmic rays (nuclei, gammas, neutrinos) of very high energies (> 100 TeV)

**Track Classification:** Cosmic rays (nuclei, gammas, neutrinos) of very high energies (> 100 TeV)